What was the Paleogene latitude of the Lhasa terrane?

A reassessment of the geochronology and paleomagnetism of Linzizong volcanic rocks (Linzhou Basin, Tibet)

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Abstract

The Paleogene latitude of the Lhasa terrane (southern Tibet) can constrain the age of the onset of the India-Asia collision. Estimates for this latitude, however, vary from 5°N to 30°N, and thus here, we reassess the geochronology and paleomagnetism of Paleogene volcanic rocks from the Linzizong Group in the Linzhou Basin. The lower and upper parts of the section previously yielded particularly conflicting ages and paleolatitudes. We report consistent $^{40}$Ar/$^{39}$Ar and U-Pb zircon dates of ~52 Ma for the upper Linzizong, and $^{40}$Ar/$^{39}$Ar dates (~51 Ma) from the lower Linzizong are significantly younger than U-Pb zircon dates (64–63 Ma), suggesting that the lower Linzizong was thermally and/or chemically reset. Paleomagnetic results from 24 sites in lower Linzizong confirm a low apparent paleolatitude of ~5°N, compared to the upper part (~20°N) and to underlying Cretaceous strata (~20°N). Detailed rock magnetic analyses, end-member modeling of magnetic components, and petrography from the lower and upper Linzizong indicate widespread secondary hematite in the lower Linzizong, whereas hematite is rare in upper Linzizong. Volcanic rocks of the lower Linzizong have been hydrothermally chemically remagnetized, whereas the upper Linzizong retains a primary remanence. We suggest that remagnetization was induced by acquisition of chemical and thermoviscous remanent magnetizations such that the shallow inclinations are an artifact of a tilt correction applied to a secondary remanence in lower Linzizong. We estimate that the Paleogene latitude of Lhasa terrane was 20 ± 4°N, consistent with previous results suggesting that India-Asia collision likely took place by ~52 Ma at ~20°N.
1. Introduction

Establishing the paleolatitude histories of terranes within N-S closing oceanic basins is a common approach to date collisional events [e.g., Gilder and Courtillot, 1997], such as the one between Greater India and the Lhasa terrane that bordered the Cretaceous Neotethyan ocean. Paleomagnetic investigations of the paleolatitude of the Lhasa terrane, which formed the southern edge of the Asian continent since the Cretaceous, started in the 1980’s and early 1990’s [Zhu et al., 1981; Pozzi et al., 1982; Westphal and Pozzi, 1983; Achache et al., 1984; Lin and Watts, 1988; Chen et al., 1993]. Recently, several groups attempted to improve on these pioneering studies by generating larger datasets and using more sophisticated analytical procedures [Chen et al., 2010; Dupont-Nivet et al., 2010b; Liebke et al., 2010; Sun et al., 2010; Tan et al., 2010; Chen et al., 2012; Liebke et al., 2012; Sun et al., 2012; van Hinsbergen et al., 2012; Huang et al., 2013; Chen et al., 2014; Lippert et al., 2014; Yang et al., 2014; Huang et al., 2015b]. Most of the later studies focused on the regionally extensive Linzizong Group composed of ~69-44 Ma lavas, ignimbrites, ash fall tuffs and sedimentary rocks. Recent paleolatitude estimates for the Lhasa terrane derived from Linzizong Group rocks vary from 5°N to 30°N, however. This range of paleolatitudes, often from the same rock units, corresponds to a discrepancy of more than 2500 km for the position of the southern Asian margin, and as a consequence, the paleomagnetically inferred age of the collision between the leading edge of Greater India and the Lhasa terrane ranges from 35 Ma to 65 Ma.

The Linzizong Group exposed in the Linzhou basin of southern Tibet (Figure 1) has been a primary target for paleolatitude studies because of its proximity to Lhasa and outstanding outcrop exposures. Highly contrasting paleomagnetic results from different intervals of the
Linzizong Group near Linzhou have been reported by Achache et al. [1984], Chen et al. [2010], Dupont-Nivet et al. [2010b], Liebke et al. [2010], Tan et al. [2010], Liebke et al. [2012], Huang et al. [2013] and Chen et al. [2014]. The source of this disparity remains unclear, although our recent work on the Linzizong Group in the Linzhou basin suggests some potential explanations.

We have shown that sedimentary rocks from the uppermost interval of the Linzizong Group (~52 Ma) have a strong bedding-parallel fabric and magnetic directional distribution that are consistent with sedimentary inclination shallowing [Huang et al., 2013]. Tan et al. [2010] reported similar phenomena and biasing in Cretaceous sediments below the Linzizong Group in this basin. We also showed that undersampling secular variation in studies focused on the volcanic intervals is another strong bias in many of the published paleomagnetic poles from the Linzizong Group [Lippert et al., 2014]. Large datasets from volcanic rocks [Dupont-Nivet et al., 2010b; van Hinsbergen et al., 2012; Lippert et al., 2014], as well as inclination shallowing-corrected results from sedimentary rocks directly overlying these volcanic rocks [Huang et al., 2013], provide a consistent Paleogene latitude estimate of ~20 ± 4°N.

Published paleomagnetic results from the lower Linzizong Group in the Linzhou basin, however, were interpreted to have been deposited at a much lower paleolatitude of ~5°N [Chen et al., 2010; Chen et al., 2014]. Such a scenario requires a > 1500 km northward shift of the Lhasa terrane within ~10 Myr relative to an essentially latitudinally stationary Eurasia. This paleogeography requires a remarkably fast northward motion of Tibetan lithosphere with a velocity of ~15 cm/yr, which is much higher than reconstructed Cenozoic intra-Asian shortening rates of ~1-1.5 cm/yr and which is almost equal to the India-Asia convergence rate in this time period [van Hinsbergen et al., 2011a; van Hinsbergen et al., 2011b]. Moreover,
such a scenario also requires that the Lhasa terrane has moved >1000 km south relative to Eurasia (requiring major extension in Asia) between the Late Cretaceous and Paleocene [Tan et al., 2010; Chen et al., 2012; Sun et al., 2012; van Hinsbergen et al., 2012; Chen et al., 2014; Lippert et al., 2014; Ma et al., 2014; Yang et al., 2014]. These scenarios of rapid yo-yo-like plate tectonic motions of the Lhasa terrane are neither predicted by the apparent polar wander path (APWP) of Eurasia since Cretaceous [Torsvik et al., 2008; Torsvik et al., 2012; Torsvik and Cocks, 2013] nor are supported by geological observations and reconstructions [Dewey et al., 1989; Yin and Harrison, 2000; Johnson, 2002; Replumaz and Tapponnier, 2003; van Hinsbergen et al., 2011b].

We recently reinvestigated a section of the lower Linzizong volcanic-bearing strata in the Nanmulin basin, 50 km northeast of Xigaze (Figure 1a), where some of these low paleolatitudes have been reported [Chen et al., 2010]. Our rock magnetic analyses, negative fold tests, and petrographic observations demonstrate that these rocks do not retain their primary magnetization, but are instead remagnetized to varying amounts. The failure to recognize this remagnetization leads to an interpreted paleolatitude of ~7°N, instead of the correct paleolatitude of ~20°N [Huang et al., 2015b]. The pervasive yet easily overlooked remagnetization of the Linzizong Group in the Nanmulin basin therefore demands a thorough rock magnetic re-investigation of the Linzizong stratigraphy in the Linzhou basin to assess if the paleolatitude estimates, including our own, calculated from these rocks are also biased by remagnetization. In the Linzhou basin, however, the structure precludes a paleomagnetic fold test and reversed directions are too sparse for a statistically meaningful reversals test; therefore, we must thoroughly characterize the magnetic petrology and thermal histories of
these units.

In this paper, we reevaluate the paleomagnetism of the lower Linzizong Group in the Linzhou basin using several approaches 1) $^{40}$Ar/$^{39}$Ar and U-Pb dates were obtained from the same samples collected throughout the entire volcanic succession for the first time. 2) We test whether the published shallow paleomagnetic directions from the lower Linzizong Group are reproducible. 3) We investigate the anisotropy of magnetic susceptibility of the lower Linzizong Group to compare with the upper intervals and to estimate the potential effect of magnetic fabrics on remanence directions [Huang et al., 2013]. 4) We provide detailed rock magnetic analyses focusing on the lower Linzizong Group, utilizing thermomagnetic experiments, hysteresis measurements, first order reversal curves (FORC) diagrams, and isothermal remanent magnetization (IRM) component analysis to evaluate the relative age and carriers of the remanent magnetization(s). Equivalent rock magnetic studies have been applied to the upper Linzizong Group [Dupont-Nivet et al., 2010b; Huang et al., 2013]. 5) To enable further comparison of volcanic rocks between the lower and upper Linzizong Group, we applied end-member modeling of acquisition curve of isothermal remanent magnetization (IRM) [Gong et al., 2009b] to identify potential differences in mechanisms for remanence acquisition. 6) Finally, we use optical and scanning electron microscopy to visually characterize the rock magnetic minerals, their relative ages of formation, and to verify rock magnetic results.
We use the described observations to propose possible mechanisms and ages of the remanence acquisition for the lower and upper volcanic rocks in the Linzhou basin and to reevaluate the reliability of published paleomagnetic results. We also discuss potential pitfalls in determining a reliable paleolatitude from cryptically remagnetized lavas. Finally, we provide an assessment of the most reliable estimates for the Paleogene latitude of the Lhasa terrane.

2. Geologic Background

The Lhasa terrane was the southernmost margin of continental Eurasia, separated from Greater India by the Neotethyan ocean prior to the India-Asia collision that began during the early Paleogene [Tapponnier et al., 1981; Burg et al., 1983; Allegre et al., 1984; Burg and Chen, 1984; Sengor, 1984; Dewey et al., 1988; Yin and Harrison, 2000]. To the south, it is separated from the Tibetan Himalaya and India affinity rocks by the Yarlung-Tsangpo suture zone. To the north, it is separated from the Qiangtang continental block by the Late Jurassic-Cretaceous Bangong-Nujiang suture zone in central Tibet (Figure 1a) [Dewey et al., 1988; Yin and Harrison, 2000; Kapp et al., 2007a]. Northward subduction of the Neotethyan oceanic lithosphere beneath Eurasia produced a Cordilleran-type continental margin, manifested by the Gangdese batholith and associated Linzizong Group volcanic rocks along the southern margin of the Lhasa terrane [Coulon et al., 1986; Yuquans, 1995; Ding et al., 2003; Lee et al., 2009; Jiang et al., 2014; Pan et al., 2014]. The Linzizong Group volcanic rocks are widely distributed in an east-west trending elongated belt along the northern edge of the Gangdese batholith (Figure 1a). The Linzizong Group erupted from ~69 to ~41 Ma determined by $^{40}$Ar/$^{39}$Ar and U-Pb dates [Coulon et al., 1986; Zhou et al., 2004; Wu et al.,
This eruptive interval spans the India-Asia collision age of ~58-50 Ma inferred from stratigraphic and metamorphic studies [Garzanti et al., 1987; de Sigoyer et al., 2000; Leech et al., 2005; Green et al., 2008; Guillot et al., 2008; Najman et al., 2010; Hu et al., 2012; DeCelles et al., 2014; Orme et al., 2014]. Thus, the Linzizong Group provides an ideal target for paleomagnetic investigations of the latitude history of the Lhasa terrane before and during the collision between the Tibetan Himalaya and the Lhasa block.

The Linzizong Group unconformably overlies strongly deformed Mesozoic strata. It is particularly well exposed and studied in the Linzhou basin, where it has a total thickness of ~3500 m and is divided into three formations, which are from bottom to top: the Dianzhong Formation (E\textsubscript{1d}), the Nianbo Formation (E\textsubscript{2n}), and the Pana Formation (E\textsubscript{2p}) (Figure 1b) [Burg et al., 1983; Xu et al., 1985; Coulon et al., 1986; TRGST, 1990; Liu, 1993; Mo et al., 2003; He et al., 2007; Leier et al., 2007; Mo et al., 2007]. E\textsubscript{1d} strata lie unconformably on the Cretaceous Takena Formation. They are composed of andesitic lavas, dacitic to rhyolitic ash-tuff layers and ignimbrite/pyroclastic flows (Figures 2a-i). A gently-dipping angular unconformity separates E\textsubscript{1d} from E\textsubscript{2n}, with bedding of E\textsubscript{2n} ~5° steeper than that of E\textsubscript{1d} (Figures 2j, k). The lower E\textsubscript{2n} consists of two main repeated sequences of mudstone, conglomerate and limestone layers at the bottom (Figure 2l). The top of E\textsubscript{2n} contains tuffs and thin-layered calcareous mudstone intercalated with thick ash deposits. E\textsubscript{2p} overlies E\textsubscript{2n} with a slightly angular unconformity and is dominated by brown-gray dacitic lapilli tuffs with lapilli in the lower part (E\textsubscript{2p}\textsuperscript{1} in Figure 1b) and interbedded sandstone, siltstone, mudstone, and ash/tuff deposits in the upper part (E\textsubscript{2p}\textsuperscript{2} in Figure 1b). Dike swarms dated to ~52 Ma are...
prevailing in the underlying Cretaceous Takena Formation, E1d, and E2n (Figures 2h, j, k) [Yue and Ding, 2006; He et al., 2007].

Several radiometric studies from the Linzhou basin, based primarily on U-Pb and $^{40}$Ar/$^{39}$Ar techniques, have yielded volcanic emplacement ages from 69 Ma to 44 Ma [BGMRXAR, 1993; Mo et al., 2003; Zhou et al., 2004; He et al., 2007; Lee et al., 2007; Ding et al., 2014]. U-Pb zircon results are generally more concordant and show less within-site scatter than whole rock $^{40}$Ar/$^{39}$Ar ages. This discrepancy may be attributed to the thermal history of the Linzhou basin [He et al., 2007], which we will discuss below. $^{40}$Ar/$^{39}$Ar dating of the dikes in E1d and E2n has yielded a plateau age of 52.9 ± 0.4 Ma [Yue and Ding, 2006], which is consistent with the U-Pb ages of the dikes intruding in the Takena Formation [He et al., 2007].

3. Paleomagnetic and Geochronologic Sampling

We have focused on volcanic rocks from E1d and E2n primarily because previous results from these units have yielded >15° discrepancies in the Paleogene latitude history for the southern Lhasa terrane [Achache et al., 1984; Chen et al., 2010; Dupont-Nivet et al., 2010b; Liebke et al., 2010; Tan et al., 2010]. To supplement the 281 characteristic remanent magnetization (ChRM) directions presented in Dupont-Nivet et al., [2010b] and Huang et al. [2013] from E2n volcanic rocks, we collected 24 volcanic sites within volcanic horizons (192 oriented cores, SH01 through SH24, site code increases up section) of E1d (Figure 1b). Standard 2.5 cm diameter paleomagnetic cores were collected using a portable gasoline-powered drill and oriented with magnetic and sun compasses (average absolute difference between magnetic and sun compass readings is 0 at all localities). Each
paleomagnetic sampling site consisted of eight independent cores collected across the thickness of the lavas and across 2-4 meters of lateral outcrop. Bedding attitudes determined by planar orientations of the top surface of the volcanic beds were measured throughout the section at several sampling sites. For the lower part of E_{1d} (SH01-SH13), a mean bedding orientation (dip azimuth = 9.2°N, dip = 16.6°; α_{95} = 4.2°) from 10 measurement sites was applied to correct the samples for bedding tilt (Table S1). For the upper part of E_{1d} (SH14-SH24), 9 measurement sites (Table S1) yielded a different mean bedding orientation (dip azimuth = 339°N, dip = 15.2°; α_{95} = 3.8°).

To detect potential thermal disturbance and to constrain the relative ages of remanence acquisition through the entire succession, samples for geochronologic studies were collected from E_{1d} (at the paleomagnetic sampling locations of site SH01) and from the tuff layers of E_{2p} (paleomagnetic Sites GS1-4 and GS124) (Figure 1b). Another sample (LD8) was collected from the lavas near the top of E_{1d} and a few hundred meters from the sampled section (Figure 1b). One additional sample (LD9) was collected from a dyke swarm intruding the Cretaceous redbeds and E_{1d} to the east of the paleomagnetic sampling locations (Figure 1b); these dykes are described by Liebke et al. [2010] and Liebke et al. [2012].

4. Geochronology

Both ^{40}\text{Ar}/^{39}\text{Ar} whole rock and U-Pb zircon geochronology were conducted on GS1-4, GS124, and SH01. Samples LD8 and LD9 were dated only using U-Pb zircon geochronology. Methods and procedures for ^{40}\text{Ar}/^{39}\text{Ar} whole rock geochronology and U-Pb zircon geochronology are described in Supplement A. The analytical data are reported in Table S2 and Table S3.
4.1 \(^{40}\text{Ar}/^{39}\text{Ar}\) Whole Rock Geochronology

The apparent age spectra and inverse isochron age results are displayed in Figures 3a-c and the results are listed in Table S2. Most samples show varying degrees of disturbance with high scatter (MSWD (Mean square weighted deviation) >> critical values, Figures 3a-c), so that only one true weighted mean plateau age could be calculated (Figures 3a-c). The remaining ages are reported by their weighted means from the spectrum analysis [Renne et al., 1998]. The number of steps selected in the age calculations does not change the reported ages significantly. Due to the high scatter, caution must be taken when interpreting the ages; however, we may use them for comparative purposes, for example between individual samples, as well as between geochronologic methods (i.e., Ar-Ar and U-Pb).

Samples GS1-4 and GS124 from E\(_{2P}\) yield dates of 50.1 ± 1.2 Ma and 50.2 ± 1.0 Ma (Figures 3a, b), respectively; these dates overlap with each other at the 95% confidence level. Sample SH01 from the older E\(_{1d}\) strata yield a similar age of 51.3 ± 0.9 Ma (Figure 3c).

4.2 U-Pb Zircon Geochronology

Zircons from samples GS1-4 and GS124 are euhedral and show clear oscillatory zoning with rare inherited cores. For sample GS1-4, thirteen spot analyses of 13 individual zircon grains provide a weighted mean age of 52.5 ± 1.2 Ma (2\(\sigma\), MSWD = 0.2), with individual zircon ages that range between 50.2 and 54.6 Ma (Figure 3a). The ages are all concordant. For sample GS124, twenty-one spot analyses of 21 individual zircon grains provide a weighted mean age of 52.0 ± 0.7 Ma (2\(\sigma\), MSWD = 0.6), with individual zircon ages that range between 50.6 and 54.4 Ma (Figure 3b). The ages are all concordant.

Samples SH01, LD8 and LD9 had low zircon recovery, but the zircons are usually...
euhedral. Most of the zircons show clear oscillatory zoning without inherited cores, but inherited cores can be found in some zircons. For sample SH01, seven spot analyses of 4 individual zircon grains (combined rim and core analyses of some individual grains, same for below) provide a weighted mean age of 63.9 ± 0.7 Ma (2σ, MSWD = 1.0), with individual zircon dates ranging from 61.4 to 65.0 Ma (Figure 3c). The ages are all concordant. Moreover, three spot analyses of two zircon grains yield concordant Proterozoic ages, and one zircon yields a concordant Neogene age (Figure 3d; Table S3). For sample LD8, nine spot analyses of five individual zircon grains show a weighted mean age of 62.9 ± 1.5 Ma (2σ, MSWD = 0.7), with individual zircon ages between 60.2 and 65.9 Ma (Figure 3e). Three spot analyses of two zircon grains also yield concordant Proterozoic ages (Table S3). For sample LD9, however, five spot analyses of 3 individual zircon grains give a Neogene weighted mean age of 12.6 ± 0.8 Ma (2σ, MSWD = 0.2), with individual zircon ages ranging from 12.4 to 13.4 Ma (Figure 3f). Concordant old ages > 400 Ma are also found (Table S3). These old ages (> 400 Ma) probably represent the recycling of old continental crust during the formation of E_{1d} volcanic rocks. Both samples SH01 and LD9 were collected within the lower part of E_{1d}, and together with the presence of old inherited cores, we believe that the young dates are meaningful and represent Neogene intrusions that cut the lower part of E_{1d}.

Comparison of the $^{40}$Ar/$^{39}$Ar ages of samples GS1-4 and GS124 to the U-Pb zircon ages of the same samples shows that the $^{40}$Ar/$^{39}$Ar ages and U-Pb zircon ages of these samples are consistent within error. By combining our results with other published radiometric ages from the Linzhou basin [BGMRXAR, 1993; Zhou et al., 2004; He et al., 2007; Ding et al., 2014], we conclude that the E_{2p} volcanic rocks were erupted between about 54 and 48 Ma. In
contrast, the $^{40}\text{Ar}/^{39}\text{Ar}$ age of samples SH01 from $E_{1d}$ is more than 10 Myr younger than the corresponding U-Pb zircon age; it is almost indistinguishable from the eruption ages of the overlying $E_{2p}$ volcanic rocks. We therefore argue that the $^{40}\text{Ar}/^{39}\text{Ar}$ system of the $E_{1d}$ volcanic rocks was disturbed by post-emplacement thermal events. The U-Pb zircon ages in our study confirms that $E_{1d}$ volcanic rocks erupted between ~68 and 62 Ma [Zhou et al., 2004; He et al., 2007; Lee et al., 2007]. We also note the presence of Neogene intrusions (~13 Ma), which are widespread in the Lhasa terrane [e.g., Coulon et al., 1986; Ding et al., 2003], but which have never been reported in the Linzhou basin.

5. Paleomagnetism

5.1 Demagnetization

We isolated ChRM directions using both thermal and alternating field (AF) demagnetization techniques. Specimens were heated and cooled in a magnetically shielded ASC oven (Model TD48-SC) that has a residual field of < 10 nT. The natural remanent magnetization (NRM) was measured on a horizontal 2G Enterprises DC SQUID magnetometer (noise level $3 \times 10^{-12} \text{Am}^2$). Samples were progressively demagnetized by temperature steps at 100, 200, 250, 300, 350, 400, 450, 500, 525, 540, 550, 560, 570, 575, 580, 600, 620, 640, 650, 660, 670 and 680°C. AF demagnetizations were applied with an in-house-developed robotized sample handler [Mullender et al., 2005] attached to a horizontal pass-through 2G Enterprises DC SQUID magnetometer (noise level $1 \times 10^{-12} \text{Am}^2$) hosted in the magnetically shielded room (residual field < 200 nT) at the Fort Hoofddijk Paleomagnetic Laboratory, Utrecht University (Netherlands). AF treatment consisted of successive steps of 5, 10, 15, 20, 25, 30, 35, 40, 45, 50, 55, 60, 65, 70, 80, and 90 mT.
For samples from E1d, three NRM components were isolated by thermal demagnetization. The first component was commonly removed below 350°C (Figure 4). After removing this low temperature component (LTC), an intermediate temperature component (ITC) was usually unblocked by 560-570°C, and a high temperature component (HTC) was unblocked at 670-680°C (Figure 4). The ITCs and HTCs share a similar ChRM direction. We note that not every specimen recorded each of these three components. For some samples, only the first component was isolated, in some samples the first and second were isolated, and in other samples all the three components were isolated (Figure 4). For most samples, AF treatment was ineffective at isolating the ChRM. Only a few samples processed by AF treatment record a similar demagnetization trajectory as observed with thermal treatment, but this direction did not decay toward the origin of orthogonal vector plots (Figure 4m).

5.2 ChRM Directions

Principal component analysis [Kirschvink, 1980] on at least five successive steps resulted in precisely determined ChRM directions for most specimens (Table S4). All ChRM directions are calculated from the thermal demagnetization results because the AF treatment was ineffective at thoroughly demagnetizing the samples. Directions with mean angular deviations (MAD) > 5º were systematically rejected from further analysis (Table S4). Lava site-mean directions with k < 50 and n < 5 were systematically discarded following volcanic data selection criteria of Johnson et al. [2008] and Biggin et al. [2008]. These criteria are consistent with the selection criteria detailed in Lippert et al. [2014] (Table S5). Fisher [1953] statistics were used to evaluate the Virtual Geomagnetic Poles (VGPs) computed from the interpreted ChRM directions; site mean VGPs located more than 45º from the formation mean VGP were
excluded from final direction calculations following the rationale of Johnson et al. [2008].

We isolated a LTC in 101 cores from all 24 sites, which together provided a mean direction of $D_g = 0.8^\circ, I_g = 41.1^\circ$ (n = 81, k = 11.7, $a_{95} = 4.8^\circ$) after applying $45^\circ$ cut-off to these individual directions (Figure 5a). The ITC and HTC are often similar to each other. The overall mean ChRM direction of the 13 sites from $E_{1d}$ that pass our data selection criteria is $D_g = 195.9^\circ, I_g = -25.4^\circ$ before tilt correction and $D_s = 193.5^\circ, I_s = -11.6^\circ$ (k = 35, $a_{95} = 7.1^\circ, \Delta D = 6.1^\circ, \Delta I = 11.8^\circ$) (Figures 5b, c; Table S5).

These ChRM directions are statistically indistinguishable from the 13 sites passing our selection criteria from $E_{1d}$ that were reported by Chen et al. [2010] and Chen et al. [2014] (xl40-49, xl68-70 and xl145-156) (Figure 5c). This follows from a positive common true mean direction (CTMD) test (classification of “C” [McFadden and Lowes, 1981]) between the 13 sites from this study and the 13 sites from Chen et al. [2010] and Chen et al. [2014]. We cannot evaluate whether the isolated ChRM directions from $E_{1d}$ are primary or not with standard field tests because bedding attitude variations are minimal throughout the sampled succession; a fold test applied to the directions would have little, if any, geologic rigor. Moreover, each of the ChRM directions is of exclusively reversed polarity, which precludes a paleomagnetic reversal test.

Results from the stratigraphically higher unit $E_{2p}$ reported by Dupont-Nivet et al. [2010b] and Huang et al. [2013] also do not allow us to apply paleomagnetic field tests because of uniform bedding orientations and a limited number of sites with reversed polarity. However, a comparison of the ChRM directions from the $E_{1d}$ volcanic rocks reported here with those reported from $E_{2p}$ volcanic rocks and inclination-shallowing-corrected sedimentary rocks
yields a discrepancy of ~30° in inclination [Dupont-Nivet et al., 2010b; Huang et al., 2013]. Below we explore possible causes for this large difference in inclination in rocks with ages within 10 Myr of each other. We measured the anisotropy of the magnetic susceptibility (AMS) of the sampled E1d volcanic rocks to test whether inclination shallowing could be related to magnetic fabrics as it has been shown for some tuffs [Rochette et al., 1999; Callot et al., 2001; Callot and Geoffroy, 2004; Schöbel and de Wall, 2014]. Secondly, we provide a detailed rock magnetic analysis along with petrographic investigations of E1d and E2p samples to identify potential remagnetization processes as was revealed in other Linzizong volcanic successions in the Nanmulin basin [Huang et al., 2015b].

6. Rock Magnetism

We applied a bevy of rock magnetic analyses, including AMS, thermomagnetic experiments, hysteresis measurements, FORC diagrams, and component analysis of the isothermal remanent magnetization (IRM) [Kruiver et al., 2001] to identify the magnetic fabric and the remanence carrier(s) of the E1d volcanic rocks. Analytical methods can be found in Supplement B.

6.1 Anisotropy of Magnetic Susceptibility

The AMS of 138 volcanic specimens from E1d were measured. The susceptibility of the E1d volcanic rocks varies considerably from $3.6 \times 10^{-5}$ SI to $7.6 \times 10^{-2}$ SI (Table S6). The degree of anisotropy (P value) is low, ranging from 1.002 to 1.048, with a mean of 1.017 (Table S6). The lavas do not exhibit uniformly oriented characteristic lineation (L) or foliation (F) (Figure 6a), and the magnetic fabric is neither dominantly oblate nor prolate (Figure 6b). The samples’ three principal ellipsoid axes ($K_{\text{max}}$, $K_{\text{int}}$ and $K_{\text{min}}$) are distributed randomly in the stereoplots.
(Figures 6c, d), confirming that no discernible fabric can be identified at the site level. These characteristics are very similar to those from our AMS studies of E$_{2p}$ volcanic rocks, but are very different from those of E$_{2p}$ sedimentary rocks [Huang et al., 2013], as well as some pristine igneous rocks (e.g., basalt) that record a magnetic fabric possibly induced by flow dynamics during the cooling of magma and secondary processes such as mineral growth during hydrothermal alteration and tectonic fracturing or deformation [MacDonadl and Palmer, 1990; Hargraves et al., 1991; Just et al., 2004; de Wall et al., 2010; Schöbel and de Wall, 2014]. We conclude that it is very unlikely that the paleomagnetic directions isolated from both E$_{1d}$ and E$_{2p}$ volcanic rocks are influenced by a preferred orientation of the magnetic minerals.

6.2 Thermomagnetic runs

Twenty representative samples of different rock types representing the entire E$_{1d}$ volcanic section were subjected to high-field thermomagnetic experiments. Typical results for volcanic samples from E$_{1d}$ show a rapid decrease in saturation magnetization up to ~350°C, ~580°C and/or ~680°C (Figures 7a-c). In general, three magnetic mineral phases corresponding to the three NRM components isolated during thermal demagnetization can be distinguished by clearly steeper decreases in magnetization slope of the thermomagnetic runs: LTC (~250-350°C), ITC (~525-580°C), and HTC (~620-680°C) (Figures 7a-c). The contribution of each component to the magnetization varies from sample to sample. The observed decreases around 350°C and around 580°C in magnetization for samples from the lower part of E$_{1d}$ (SH9.7) define the LTC and ITC (Figure 7a). Samples from the upper part of E$_{1d}$ (SH18.8, and SH24.7) are characterized by the appearance of the HTC in addition to the LTC and ITC (Figures 7b, c). We interpret the LTC to be associated with Ti-rich titanomagnetite, the ITC to
be Ti-poor titanomagnetite, and the HTC is interpreted to be carried by hematite following Dunlop and Ödemir [1997].

6.3 Hysteresis Loops, IRM & Back-field Curves, and FORC Diagrams

Hysteresis measurements were applied to thirteen representative samples of different lithologies. We observed two distinctive types of hysteresis loops for samples from E1d. Type 1 loops are narrow and are essentially saturated at 600 mT (Figure 7d). The corresponding IRM curves are also completely saturated by 600 mT (Figure 7d). Bulk coercivity ($B_c$) values and remanent coercivity ($B_{cr}$) values of the samples are usually low (<60 mT) (Table S7). Type 2 loops are generally wasp-waisted (Figures 7e, f), reflecting the presence of two or more magnetic components with strongly contrasting coercivities [Roberts et al., 1995; Tauxe et al., 1996]. IRM acquisition curves remain unsaturated up to 2 T, indicating the presence of high coercivity minerals (Figures 7e, f). Both $B_c$ values and $B_{cr}$ values are high, ranging from ~44 to 330 mT and from ~85 to 560 mT, respectively (Table S7).

Similar to our observations of the hysteresis loops, we distinguish two distinct groups of FORC diagrams. The FORC diagrams for samples with type 1 loops are characterized by contours that diverge away from the origin and are spread along the $H_u$ axis (Figure 7g). This contour pattern is consistent with the typical behavior of interacting pseudo-single domain (PSD) to multiple domain (MD) magnetic particles [Pike et al., 1999; Roberts et al., 2000]. The FORC diagrams of samples with type 2 loops are characterized by the appearance of two contributions with different coercivities (Figures 7h, i). A dominant population of SD or PSD grains with very low coercivity is apparent along $H_u = 0$ mT is shown, but an additional contribution with distinctly higher coercivities and negligible magnetic interaction is also
clearly displayed. We interpret the SD and PSD-MD magnetic particles with low to intermediate coercivities to be titanomagnetite, whereas magnetic particles with high coercivity are inferred to be hematite. These interpretations are consistent with the thermomagnetic data described above.

6.4 IRM Component Analysis

In general, the IRM acquisition curves are preferably fit by four IRM components: component L1 with $B_{1/2}$ (the field at which half of SIRM is reached) of ~10 mT and dispersion parameter (DP) of ~0.35 (log units); a soft component L2 with $B_{1/2}$ ranging from 25 mT to 35 mT and DP varying from 0.22 to 0.35; a relatively hard component L3 with higher $B_{1/2}$ (50-150 mT) and notably variable DP (0.14-0.45); and a much harder component L4 with $B_{1/2}$ more than 400 mT and a DP of ~0.25 (Figures 7j-l; Table S8). Component L1 with very low coercivity constitutes only a few percent of the SIRM; it is only required to fit the skewed-to-the-left distribution of component L2 and is not given physical meaning other than being the result of thermally activated component L2 particles [Egli, 2004; Heslop et al., 2004].

We interpret component L2 and component L3 to represent Ti-poor titanomagnetite and Ti-rich titanomagnetite, respectively. These two components carry ~75%-100% of the SIRM for most samples in the lower part of E1d (Table S8). The high dispersion parameter of these components suggests a broad magnetic grain-size distribution. We interpret component L4, with its very high coercivity, to be hematite. It contributes up to 96% of the SIRM for samples in the upper part of E1d (Table S8).

These interpretations of the IRM component analysis are consistent with results from our other rock magnetic experiments. Ti-poor and Ti-rich titanomagnetite are the dominant
magnetic carriers for samples from the lower part of E₁d. In contrast, samples from the upper part of E₁d are characterized by large amounts of hematite contributing to the SIRM.

7. End-Member Modeling of Magnetic Components

To illustrate the mechanism of remanence acquisition and to detect potential remagnetization of the Linzizong Group volcanic rocks from both E₁d and E₂p, we applied the method of end-member modeling of magnetic components as retrieved from IRM acquisition curves [Gong et al., 2009b]. End-member modeling can be helpful for identifying remagnetized and non-remagnetized sedimentary rocks independent of paleomagnetic field tests [Gong et al., 2009b; Van Hinsbergen et al., 2010; Meijers et al., 2011; Dekkers, 2012; Aben et al., 2014; Huang et al., 2015c; Huang et al., 2015a]. Our previous work on the Linzizong Group in the Nanmulin basin has shown that it is also a powerful tool for diagnosing remagnetization in volcanic rocks [Huang et al., 2015b].

We measured IRM acquisition curves for 123 volcanic samples covering the entire sampled E₁d interval and 153 volcanic samples covering the entire E₂p interval. Procedures of the measurement and data analysis are the same as those described in Huang et al. [2015b].

7.1 Volcanic Rocks from E₁d

For samples from E₁d, the calculated two-end-member model by the end-member solutions has a low r² value of 0.60219. The three end-member model has a r² value of 0.82727 and a convexity of -2.4634 (Figure 8a). The four-end-member model has an even higher r² of 0.87399 but a lower convexity of -2.0566 (Figure 8a). Models with five or more end-members are characterized by virtual duplication of end-members, which indicates that the dataset is over-interpreted in these scenarios. Comparison of the models with three and four
end-members shows that the end-members DEM1 (D stands for ‘Dianzhong formation’) and DEM3 are indistinguishable for both options, whereas DEM2 in the three end-member model is divided into DEM2 and DEM4 in the four end-member model (Figures 8b, c). This indicates that the four end-member model does not discriminate more distinctive and geologically interpretable end-members than the three end-member model. Moreover, a break in slope is observed at three end members in the plot of $r^2$ vs. the number of end members (Figure 8a). We therefore use the three end-member model to interpret our dataset from the $E_{1d}$ volcanic rocks.

We applied IRM component analysis to these three end-members [Kruiver et al., 2001]. Four components (components D1 through D4, increasing from soft to hard magnetically) are required to fit the normalized IRM-acquisition curves of the three end members. DEM1 can be fit with two components (component D2 and component D4, Figure 8e). Component D2 with $B_{1/2}$ of ~30 mT typically represents Ti-poor titanomagnetite [Day et al., 1977; Lowrie, 1990]. The contribution of this component to SIRM is very low (~3%). Component D4 is probably hematite as indicated by the high $B_{1/2}$ of ~420 mT [Kruiver and Passier, 2001]. It contributes more than 95% to the SIRM (Table S8). The contribution of D2 to SIRM is very low (~3%). End-member DEM2 can be fit with component D3. We interpret it to be Ti-rich titanomagnetite, because of the fairly high $B_{1/2}$ of ~90 mT [Day et al., 1977; Lowrie, 1990] (Figure 8f). The IRM acquisition curve for DEM3 requires three IRM components for a good fit: component D1, a very low-coercivity component, constitutes ~10% of the SIRM; it is only required to fit the skewed-to-the-left distribution of component D2 and is not given physical meaning other than being the result of thermally activated component D2 particles [Egli, 2004; Heslop et al., 2004]. Component D2 and component D4 typically represent Ti-poor titanomagnetite and hematite,
respectively [Day et al., 1977; Lowrie, 1990; Kruiver and Passier, 2001]. The contribution of component D2 to the SIRM is up to 85% and component D4 has a limited contribution to the SIRM of 4-6% (Figure 8g; Table S8). Inspection of the IRM component analysis of the end-members shows that the four components required to fit them are indistinguishable from those components for the IRM acquisition curve fitting described in our rock magnetic studies above. This implies that the three end-members are geologically interpretable [Weltje, 1997].

Most samples from E1d are mixtures of the three end members in the ternary plot (Figure 8d). However, the contribution of DEM1 varies widely and the contribution of DEM3 is usually limited (Figure 8d). There are also several samples that plot on the baseline between DEM2 and DEM3. Together with the above analysis for the mineralogic interpretation of the three end-members, we conclude that magnetic carriers of the volcanic rocks from E1d have variable contributions from hematite. Notably, the fraction of the remanence carried by hematite, as distinguished by end-member modeling, increases up-section toward the unconformity with E2n.

7.2 Volcanic Rocks from E2p

Similar to our observations from the volcanic rocks from E1d, a three-end member model (PEM1 to PEM3, P stands for ‘Pana formation’) is also preferred for the interpretation of the dataset from E2p volcanic rocks (Figure 9), paleomagnetic results from which were previously reported by Dupont-Nivet et al. [2010b] and Tan et al. [2010]. The end-members of these rocks also require four IRM components (components P1 through P4, increasing from magnetically soft to hard) for a good fit of the normalized IRM-acquisition curves of the three end members. Components P1 - P4 have the same physical meaning as components D1 - D4 (Figures 9e-g;
Table S8). PEM1 represents a combination of hematite (contribution to SIRM up to ~90%) and Ti-poor titanomagnetite (contribution to SIRM ~10%, higher than that in DEM1) (Table S8). PEM2 is interpreted to be Ti-rich titanomagnetite. PEM3 is composed of Ti-poor titanomagnetite (contribution to SIRM approaching ~90%) and minor contributions of hematite (contribution to SIRM around 5%).

In a ternary plot, most samples from E$_{2p}$ are mixtures of PEM2 and PEM3 with minor contributions from PEM1 (Figure 9d). Moreover, careful inspection of the ternary plot reveals that more samples plot closer to PEM2 than PEM3, which indicates that the contribution of PEM2 is higher than that of PEM3 for most samples. Together with our above analysis that Ti-poor titanomagnetite also contributes ~10% to the SIRM in PEM1, we thus argue that titanomagnetite is the dominant magnetic carrier of the volcanic rocks from E$_{2p}$, whereas the contribution from hematite is minor. This conclusion is consistent with previously reported rock magnetic results from E$_{2p}$ volcanic rocks [Dupont-Nivet et al., 2010b; Huang et al., 2013].

8. Petrography

We analyzed thin sections of volcanic rocks from both E$_{1d}$ and E$_{2p}$ to identify textural relationships and diagenetic conditions of magnetic minerals with optical microscopy and scanning electron microscopy (SEM). Five representative samples from E$_{1d}$ (SH2.8, SH5.5, SH18.4, SH18.6, and SH18.7) and four samples from E$_{2p}$ (GL2.3, GL9.7, GL14.7 and GL29.4) were investigated with optical microscopy. Five representative samples from E$_{1d}$ (SH2.4, SH4.2, SH13.3, SH18.4, and SH24.8) and two samples from E$_{2p}$ (GL8.5 and GL27.3) were also analyzed on a Hitachi S3400 SEM, operated at 20 keV and 40-60 nA at the Electron Microscope Unit at the University of Hong Kong (China).
8.1 Optical Petrography

Detailed petrographic observations of thin sections reveal that samples from E$_{1d}$ are generally characterized by strong alteration, whereas the amount of alteration observed in samples from E$_{2p}$ is low. Samples from E$_{1d}$ are grey to reddish andesitic lavas. They have porphyritic texture with plagioclase/quartz lapillis distributed in a silicic matrix (Figures 10a-f). Alteration of plagioclase to sericite is prevalent (Figures 10a-f). Accessory minerals include biotite, epidote, rutile, and iron oxides. Opaque minerals (mostly hematite) are very common as accessory phases in all samples. Three distinct populations can be recognized. The first population consists of independent crystals distributed randomly within the matrix (Figures 10a-c). A second population clusters at margins of preexisting minerals, and appears to be an alteration product of biotite and feldspar (Figures 10a-c, f). A third population consists of small grains distributed at margins of the silicate minerals or along the cracks (Figures 10d, e). The first and the second populations are very common in samples SH2.8 and SH5.5 (i.e., lower in the section), whereas the third population is dominant in samples SH18.4, SH18.6, and SH18.7 (i.e., higher in the section).
Samples from $E_n$ are tuff with a pyroclastic texture characterized by scattered lapilli distributed in an ash matrix, as described in detail by Huang et al. [2013] (Figures 10g-i). Plagioclase is slightly altered to clay minerals, calcite, and chlorite. The opaque minerals (mostly (titano)magnetite) are common as an accessory phase in all the samples. They often occur as large isolated crystals distributed randomly within the ash matrix or clusters in the volcanic fragments.

8.2 SEM Observations

Additional textural information about the iron oxides is provided by SEM observations. Two types of iron oxides, titanomagnetite and hematite with contrasting petrographic features, are generally present in samples from $E_1$. However, the abundances of titanomagnetite and hematite change from sample to sample. In samples SH2.4 and SH4.2 from lower in the section, titanomagnetite grains are abundant, whereas hematite is rarely observed. Titanomagnetite grains usually occur as euhedral to subhedral grains ranging in size from a few $\mu$m up to 500 $\mu$m (Figures 11a-c). Most of the titanomagnetite grains show solid-state oxidation-exsolution features with a trellis of ilmenite distributed in the Ti-poor magnetite. These textures indicate a magmatic origin [Craig, 2001; Turner et al., 2008] (Figures 11a-c). We also observe titanomagnetite with oxidation-exsolution features in samples SH13.3, SH18.4, and SH24.8, but note that the abundance is remarkably low (Figures 11d-l). Hematite in our samples differs substantially from titanomagnetite: it has a different crystal habit and its distribution is also different. It is usually fine-grained and occurs either intergrown within titanomagnetite grains (Figures 11f, g, j, k), around the silicate minerals (Figures 11f), or along folia within the matrix and larger crystals (Figures 11d, e).
Titanomagnetite is abundant in samples from E\textsubscript{2p} (Figures 11m-o), compared to samples from the upper intervals of E\textsubscript{1d}. Titanomagnetite grains occur in a range of sizes, from less than 10 µm to over 400 µm (Figures 11m-o). They are concentrated in the lapilli (Figure 11m). The titanomagnetite grains are usually euhedral and exhibit oxidation-exsolution features. These observations suggest that the host tuff was formed during a slow-ascending eruption [Turner et al., 2008] and that the titanomagnetite is primary.

These textural observations suggest that titanomagnetite has a magmatic origin, whereas hematite is secondary and formed authigenically after volcanism. Given that SEM observations indicate that titanomagnetite is the predominant iron oxide and end-member modeling of the magnetic components indicates that this titanomagnetite carries the bulk of the magnetic remanence in samples from E\textsubscript{2p} volcanic rocks in the Linzhou basin, we conclude that, mineralogically, samples from E\textsubscript{2p} volcanic rocks show minor to no signs of chemical alteration and thus remagnetization. In contrast, petrographic observations clearly indicate that hematite is of secondary origin, and we suspect that samples from E\textsubscript{1d} with hematite as a significant remanence carrier likely records a secondary magnetization. That is, at least some, if not most, samples from E\textsubscript{1d} in the Linzhou basin might be partially remagnetized, similar to what we found in the Nanmulin basin [Huang et al., 2015b].

9. Discussion

9.1 The Case for Remagnetization

Inclination shallowing caused by sedimentary processes has been advocated to explain the shallow bias in paleomagnetic inclinations recorded by detrital sedimentary rocks [King, 1955; Lovlie and Torsvik, 1984; Anson and Kodama, 1987; Deamer and Kodama, 1990;
However, volcanic rocks are, in principle, immune to inclination shallowing biases. Unlike the strongly shallowed inclinations in the sedimentary cover of the Low Cretaceous Xigaze ophiolites [Huang et al., 2015a], the redbeds from the Upper Cretaceous Takena formation [Tan et al., 2010], and the sedimentary rocks from $E_{2p}$ in the Linzhou basin [Huang et al., 2013], the lavas of $E_{1d}$ and $E_{2p}$ do not have a characteristic magnetic anisotropy fabric that is consistent with flattening or an otherwise strong igneous fabric (Figure 6) (see also Huang et al. [2013] and Chen et al. [2014]).

Paleosecular variation (PSV) of the Earth’s magnetic field can lead to an intrinsic variation of $>30^\circ$ in inclination in spot readings of the geomagnetic field [Johnson et al., 2008; Tauxe et al., 2008; Deenen et al., 2011]. Large datasets from volcanic rocks in $E_{2p}$ with a sufficiently long range of emplacement ages display directional distributions that suggest PSV is fully characterized [Dupont-Nivet et al., 2010b; Lippert et al., 2014]. These conclusions are further supported by the similarity between ChRM directions of the lavas from $E_{2p}$ and the inclination shallowing-corrected sedimentary rocks from $E_{2p}$ [Huang et al., 2013].

The $E_{2p}$ and $E_{1d}$ volcanic deposits described here were dated using whole rock $^{40}\text{Ar}/^{39}\text{Ar}$ and zircon U-Pb geochronology from the same volcanic samples. For volcanic rocks from $E_{2p}$, the similarity between whole rock $^{40}\text{Ar}/^{39}\text{Ar}$ and zircon U-Pb ages suggests that thermal or chemical resetting is inconsequential. Our analyses based on rock magnetic investigations and petrographic observations also provide no evidence for substantial alteration and secondary magnetic mineral growth. Thus, there is no rock magnetic, mineralogic, or
geochronologic reason to doubt a primary origin of the NRM isolated from the E\textsubscript{2p} volcanic rocks. In contrast, volcanic rocks from E\textsubscript{1d} have whole rock $^{40}\text{Ar}/^{39}\text{Ar}$ ages that are statistically younger than the zircon U-Pb ages and that are of similar or even younger ages than ages of the overlying E\textsubscript{2p} volcanic rocks. These observations indicate thermal and/or chemical resetting of the E\textsubscript{1d} volcanic rocks. Together with the widespread secondary hematite growth revealed by our rock magnetic and petrographic investigations, we conclude that the NRM isolated from the E\textsubscript{1d} volcanic rocks was acquired in part during post-depositional alteration. To summarize: the E\textsubscript{1d} sequence has been remagnetized.

9.2 Potential Remagnetization Mechanisms

Acquisition of magnetization after the formation of a rock, termed remagnetization, can obscure or remove the primary magnetizations and thus limit the utility of paleomagnetism in global (plate) tectonic reconstructions [Elmore et al., 2012; Font et al., 2012]. Widespread remagnetization events are usually linked to large-scale tectonic processes [McCabe et al., 1983; Miller and Kent, 1988; Chen and Courtillot, 1989; Appel et al., 1995; Molina Garza and Zijderveld, 1996; Weil and Van der Voo, 2002; Font et al., 2006; Rapalini and Bettucci, 2008; Font et al., 2012] and can affect both sedimentary rocks [Appel et al., 1991; Gong et al., 2009a; Roberts et al., 2010; Liu et al., 2011; van der Voo and Torsvik, 2012] and igneous rocks [Harlan et al., 1996; Edel and Aifa, 2001; Geissman and Harlan, 2002; Otofuji et al., 2003; Borradaile et al., 2008; Preeden et al., 2009]. Suspected mechanisms for remagnetization include elevated temperature, chemical alteration, and secondary mineral growth. These processes can act alone or in concert, and they frequently induce changes to the NRM, to the magnetic mineralogy, or both [Jackson and Swanson-Hysell, 2012].
9.2.1 Acquisition of Thermoviscous Remanent Magnetization

The volcanic record of the southern Lhasa terrane indicates a magmatic “flare-up” at ~51 Ma [Kapp et al., 2007b; Lee et al., 2009; Jiang et al., 2014]. In the Linzhou basin, this flare-up is characterized by the eruption of large volumes of E\text{2p} volcanic rocks, intrusion of dykes into E\text{2n}, E\text{1d}, and the Cretaceous Takena formation, as well as emplacement of granitic plutons into Triassic strata [Yue and Ding, 2006; He et al., 2007]. Therefore, we must consider the regional thermal disturbance this volcanism must have provided. This thermal disturbance could have had a strong effect on volcanic rocks from E\text{1d}, given that the volcanic interval from E\text{2p} (~1900 m) is much thicker than E\text{1d} (~270 m) [He et al., 2007] and that early Eocene dykes are widespread within E\text{1d} (Figures 2h, j, k). Our geochronologic studies of the E\text{1d} lavas show a discrepancy of ~10 Myr between the U-Pb dates and \textsuperscript{40}Ar/\textsuperscript{39}Ar dates (Figure 3c). Moreover, the \textsuperscript{40}Ar/\textsuperscript{39}Ar ages of the E\text{1d} lavas are similar to the ages of the dykes and E\text{2p} tuffs, indicating that thermal disturbance of \textsuperscript{40}Ar/\textsuperscript{39}Ar system of the E\text{1d} volcanic rocks is probably induced by intrusion of the dykes and eruption of the overlying E\text{2p} volcanic rocks. Therefore, we conclude that it is very likely that the conductive cooling of dykes and thick tuffs at ~52 Ma thermally activated the titanomagnetite grains to produce a thermoviscous remanent magnetization (TVRM) in the E\text{1d} volcanic rocks.

Several studies have shown how remagnetization by TVRM has been acquired in rocks subjected to post-depositional thermal events [Kent, 1985; Kent and Miller, 1987; Smith and Verosub, 1994; Harlan et al., 1996; Dunlop et al., 1997; Valet et al., 1998; Enkin et al., 2000]. Acquisition of TVRM is based on the relationships between relaxation time and blocking
temperatures [Néel, 1949; Pullaiah et al., 1975; Middleton and Schmidt, 1982; Dunlop et al., 2000], which implies that the primary NRM can be partially reset below the Curie temperature when the temperature is elevated for a sufficiently long duration. A secondary NRM can be acquired during subsequent (very slow) cooling from these elevated temperatures.

Given closure temperatures of the K-Ar system of ~300-350°C for biotite and of ~150-300°C for potassium feldspar [McDougall and Harrison, 1999], the disturbance observed in the whole rock ⁴⁰Ar/³⁹Ar system of the lava flows from E₁d suggests that these rocks probably have been heated up to ~300°C. Assuming a minimum of 20 kyr duration for remagnetization (based on the assumption that secular variation has been adequately averaged), SD magnetite with maximum unblocking temperatures of 560-570 °C (at laboratory heating time of 45 minutes, same for the following calculations) require temperatures ≥550°C to completely reset the primary NRM [Pullaiah et al., 1975]. However, the required temperatures could be ~490 °C if the magnetite is of PSD size [Middleton and Schmidt, 1982]. The required temperature will be (much) lower if the maximum unblocking temperatures of (titano)magnetite decrease as a function of grain size or primary composition [Pullaiah et al., 1975; Middleton and Schmidt, 1982]. For example, SD titanomagnetite and PSD titanomagnetite with maximum unblocking temperatures both of 500°C can be completely reset at ~450°C and ~380°C, respectively, if the heating persists for 20 kyr. With an unblocking temperature of ~350°C, however, SD and PSD titanomagnetite can be completely reset at ~260°C and ~180°C, respectively, if the higher temperature would have prevailed for as few as 100 years. We have detected both Ti-rich titanomagnetite (unblocked at 300-350°C) and Ti-poor titanomagnetite (unblocked at 550-570°C) with domain states of
PSD-MD in our rock magnetic investigations (Figure 7). Therefore, we conclude that the primary NRM carried by Ti-rich titanomagnetite easily could have been overprinted by a secondary TVRM similar to the NRM recorded by E₂p volcanic rocks during the thermal disturbance. In contrast, only part of the primary NRM residing in Ti-poor titanomagnetite would have been overprinted by a similar TVRM. Thus, the isolated ITC carried by Ti-poor titanomagnetite is a combination of a predominantly primary NRM with reversed polarity and a minor TVRM with normal polarity, resulting in a composite direction that has a reversed polarity but a shallower inclination than the primary NRM.

Careful inspection of the isolated LTC, however, shows that its mean direction is shallower than both the NRM recorded by volcanic rocks from E₂p and the viscous overprint induced by the present day magnetic field (Figure 5a). It is actually very close to the NRM recorded by volcanic rocks from E₂p after tilt correction [Dupont-Nivet et al., 2010b]. Given that PSD titanomagnetite with maximum unblocking temperatures of 350 °C can be completely reset at ~120°C if heating persists for 100 kyr [Middleton and Schmidt, 1982], we suggest that it is likely that the TVRM recorded by Ti-rich titanomagnetite in E₁d at ~52 Ma could have been reset again whereas the NRM of the Ti-poor titanomagnetite was not significantly affected. This thermal event might be caused by Neogene intrusions in the lower part of E₁d (dated at ~13 Ma; Figures 3d, f). The acquisition of the new TVRM likely occurred after E₂p had been tilted ~30° toward the north. This new TVRM might also be carried by the dykes intruded in E₁d, because the in situ magnetization directions of the dykes are similar to that of the in situ LTC of the E₁d lavas; the magnetic carrier is Ti-rich titanomagnetite in both units [Liebke et al., 2010; Liebke et al., 2012].
9.2.2 Acquisition of Chemical Remanent Magnetization

We used the rock magnetic and petrographic studies described above to characterize the variable abundance of secondary hematite in some of the lavas from E₁d. These observations are similar to those of the partially remagnetized volcanic rocks of the Linzizong Group in the Nanmulin basin [Huang et al., 2015b]. These similarities suggest that a chemical remanent magnetization (CRM) carried by the secondary hematite is responsible for at least some of the remagnetization. Post-eruptive precipitation and growth of hematite to produce CRMs during low-temperature hydrothermal alteration of igneous rocks has been described in several settings [Edel and Schneider, 1995; Edel and Aifa, 2001; Geissman and Harlan, 2002; Ricordel et al., 2007; Preeden et al., 2009; Parcerisa et al., 2013]. For example, oxidizing hydrothermal fluids migrating along a structural or stratigraphic plane (e.g., unconformity, fault, fracture, and/or joint) can leach iron from magmatic (titano)magnetite and Fe-rich silicates in the igneous body; this iron can then precipitate as hematite at temperatures well below the Curie temperature [Elmore et al., 1993; Dekkers, 2012; Elmore et al., 2012]. The direction of the NRM carried by this secondary hematite can be very different from that of the primary remanent magnetization if this process happens after the tilting or large-scale tectonic motion of the terrane. Importantly, it does not matter whether the igneous rocks are completely or only partially remagnetized, the isolated ChRM cannot be used in a meaningful way for paleogeographic reconstructions.

Our rock magnetic studies suggest that the amount of secondary hematite decreases downward from the unconformity between E₁d and E₂n. This trend is consistent with remagnetization caused by paleoweathering and laterization of the lavas of E₁d due to
percolation of meteoric water prior to deposition of E\textsubscript{2n} [Schmidt et al., 1976; Kumar and Bhalla, 1984; Kumar, 1986; Evans et al., 2002; Théveniaut and Freyssinet, 2002]. This scenario is further supported by field observations that the lavas of E\textsubscript{1d} are unconformably overlain by red beds at the base of E\textsubscript{2n} (Figures 2j-l).

Alternatively, the remagnetization of E\textsubscript{1d} could have been caused by migration of a high-temperature magmatic hydrothermal fluid along the base of the sedimentary strata from the lowermost E\textsubscript{2n} and the unconformity between E\textsubscript{1d} and E\textsubscript{2n} during the eruption of the thick tuffs from E\textsubscript{2p} at \(\sim\)52 Ma. This event was accompanied by the intrusion of some of the dykes in E\textsubscript{1d} and E\textsubscript{2n}. Strong hydrothermal alteration of E\textsubscript{1d} lavas is observed not only near the unconformity but also close to the dykes (Figure 2g).

We cannot rule out either of the two processes described above; in fact, we suspect that they worked in concert for the acquisition of CRMs for the reasons described next. During the demagnetization of the E\textsubscript{1d} volcanic rocks, the HTC carried by hematite is of reversed polarity with a low inclination. If the hematite formed during the paleoweathering and laterization between \(~62\) Ma and \(~54\) Ma when E\textsubscript{1d} was slightly tilted to the south by about 5º (the dip of the red beds at the lowermost of E\textsubscript{2n} strata is \(~5º\) steeper than that of E\textsubscript{1d}), then this hematite probably would have acquired a CRM similar to the primary magnetization with reversed polarity (\textit{i.e.}, Chrons C27r, C26r, C25r, or C24r; \textit{n.b.:} intervals between \(~62\) Ma-54 Ma with normal magnetic polarity are very short [Gradstein et al., 2012]). This CRM should have recorded a steeper inclination than what we observe. However, if the hematite crystallized during the migration of magmatic hydrothermal fluid at \(~52\) Ma when E\textsubscript{1d} was tilted \(~15º\) toward the south (dip of E\textsubscript{2p} rocks is \(~15º\) steeper than that of E\textsubscript{1d}, Figure 1b), then we expect it
to acquire a CRM similar to the TRM of $E_{2p}$. This CRM would have normal polarity and a very high inclination before tilt correction. We observe neither of these predictions. However, if both of the suggested processes are responsible for the CRM of $E_{1d}$, then we expect ChRM directions carried by secondary hematite to be similar to the directions we observe. Therefore, we favor a scenario in which the CRMs carried by the secondary hematite were acquired during two distinct events.

In summary, we suggest that the remagnetization of lavas from $E_{1d}$ is a combination of TVRM and CRM acquisitions (Figure 12). We suggest that there were two episodes for the acquisition of TVRM (TVRM$_1$ and TVRM$_2$). In our scenario, TVRM$_1$ is carried by Ti-poor titanomagnetite and was induced by the thermal disturbance associated with the eruption of $E_{2p}$ volcanic rocks and intrusion of dykes into $E_{1d}$ (Figure 12c). TVRM$_2$ is carried by Ti-rich titanomagnetite and was acquired after $E_{2p}$ had been tilted ~30° toward the north at ~13 Ma (Figure 12c). The CRM is carried by secondary hematite, which also may have been formed during two stages. The first stage for the acquisition of CRM (CRM$_1$) occurred sometime between ~62 Ma and 54 Ma when $E_{1d}$ was tilted ~5° toward the south and $E_{2n}$ was deposited (Figure 12b). The second stage for the acquisition of CRM (CRM$_2$) probably occurred at the same time as TVRM$_1$ at ~52 Ma, during the eruption of $E_{2p}$ and intrusion of dykes throughout $E_{1d}$. The $E_{1d}$ sequence was tilted ~15° toward the south at this time (Figure 12c). We acknowledge that this sequence of events is non-unique and other remagnetization scenarios may explain the observed paleomagnetic data. Our intent is not to fully explain the events responsible for the remagnetization, but to emphasize that the preponderance of field and laboratory observations indicate that the ChRM directions recorded in $E_{1d}$ lavas are not
primary. We have illustrated one plausible sequence of events that 1) can explain the observed directions, 2) is consistent with field observations and regional geochronology, and 3) can be tested with additional geochronology, thermochronology, and structural analysis.

9.3 Paleogene latitude of the Lhasa Terrane and Implications for Constraining the Age of the India-Asia Collision

We have reassessed the paleomagnetic results of the Linzizong volcanic rocks in the Linzhou basin using detailed rock magnetic investigations, petrographic observations, and geochronologic data. Our analyses show that the discrepancy of ~30° in tilt-corrected inclinations between volcanic rocks from E$_{1d}$ and E$_{2p}$ is an artifact of an inappropriate tilt-correction applied to a secondary remanence in remagnetized rocks. Lavas from E$_{1d}$ have been remagnetized by secondary CRMs and TVRMs, whereas volcanic rocks from E$_{2p}$ do not show any evidence of secondary magnetic mineral growth or thermal resetting. Therefore, we conclude that ChRM directions from E$_{1d}$ volcanic rocks in the Linzhou basin cannot be used to calculate the Paleogene latitude of the Lhasa terrane. Volcanic rocks from E$_{2p}$ that do record a primary NRM and have a directional distribution that can be explained by paleosecular variation are interpreted to accurately record the Paleogene latitude of the Lhasa terrane. We assert that the mean inclination of 40.3 ± 4.5° (95% confidence limit) measured in the volcanic rocks of the E$_{2p}$ unit is reliable for tectonic reconstructions [Dupont-Nivet et al., 2010b]. Inclinations obtained from sedimentary rocks of the upper E$_{2p}$ unit in the Linzhou basin that have been corrected for inclination shallowing using two independent correction methods (40.0° ([33.1°, 49.5°], 95% confidence limit) and 41.3 ± 3.3° (95% confidence limit)) are
indistinguishable from the primary E$_{2p}$ volcanic rocks directions [Huang et al., 2013]. We also note that our recent results from ~52 Ma volcanic rocks in the Nanmulin basin ~200 km to the west of the Linzhou basin suggest a primary inclination of 38.1° ([35.7°, 40.5°], 95% confidence limit) [Huang et al., 2015b]. Collectively, these paleomagnetic results from the Linzizong Group indicate that the southern Lhasa terrane was located at ~20 ± 4°N (assuming a 100 percent geocentric axial dipole field) during early Eocene time [BGMRXAR, 1993; Mo et al., 2003; Zhou et al., 2004; He et al., 2007; Huang et al., 2015b].

Given that our results show that E$_{1d}$ volcanic rocks in the Linzhou basin have been remagnetized, we conclude that the results from E$_{1d}$ volcanic rocks reported by Chen et al. [2010] and Chen et al. [2014] are also likely to be remagnetized and therefore cannot be used for paleolatitude calculations. Our results preclude the large and rapid Late Cretaceous to early Paleogene latitude oscillation of the Lhasa terrane inferred by Chen et al. [2014]. Instead, the conclusions presented here are consistent with recent reviews of robust paleomagnetic data from Cretaceous and younger volcanic rocks and other inclination shallowing-corrected sedimentary rocks from the Lhasa terrane that demonstrate relatively stable paleolatitudes of Lhasa terrane from ~110 to 50 Ma [van Hinsbergen et al., 2012; Lippert et al., 2014]. Our results are also consistent with the predictions from the Eurasian apparent polar wander path that suggest no significant paleolatitudinal motions for stable Eurasia occurred during Late Cretaceous to early Cenozoic times [Torsvik et al., 2012], and the finding that not more than 600-750 km of intra-Asian shortening has occurred during the Cenozoic [Dewey et al., 1989; Lippert et al., 2011; van Hinsbergen et al., 2011b].

We conclude that the NRM carried by Ti-rich titanomagnetite in the dykes and E$_{1d}$ lavas
of the lower Linzizong volcanic rocks in the Linzhou basin was overprinted by later thermal resetting. We also have shown that lavas and sedimentary rocks of the lower Linzizong Group have been variably remagnetized by low-temperature fluid flow and precipitation of secondary hematite in the Nanmulin basin [Huang et al., 2015b] and by high-temperature fluid alteration and thermal resetting in the Linzhou basin. Our work on the remagnetized lower Linzizong volcanic rocks in the Linzhou basin also highlights that hematite in volcanic rocks should be suspected of carrying a secondary, and sometimes misinterpreted, remanent magnetization. We implore future paleomagnetic studies in this region to be vigilant in appraising the remanence carriers and thus inferring primary magnetizations. Our work underscores the utility of acquiring thorough rock magnetic, petrographic, and geochronologic information in tectonically-directed paleomagnetic research. Paleomagnetic studies that have characterized paleosecular variation, assessed datasets for sedimentary inclination shallowing, and tested for remagnetization yield a consistent Paleogene latitude of ~20 ± 4°N for the Lhasa terrane. This paleolatitude indicates that the collision between the Tibetan Himalaya and the Lhasa terrane began by ~52 Ma, if not earlier [Dupont-Nivet et al., 2010b; Najman et al., 2010; van Hinsbergen et al., 2012; DeCelles et al., 2014; Garzanti and Hu, 2014; Hu et al., 2014; Lippert et al., 2014].

10. Conclusions

Paleomagnetic studies of the Paleogene Linzizong Group in the Linzhou basin by several independent teams have yielded a range of paleolatitude estimates for the Lhasa terrane from ~20°N from the volcanic rocks in upper intervals (E2p) to ~5°N from the volcanic rocks in the lower intervals (E1d) [Chen et al., 2010; Dupont-Nivet et al., 2010b; Tan et al., 2010; Chen et
Here, we have reevaluated this section by thorough sampling of the volcanic rocks from both the lower and upper intervals to test the reproducibility of the previous results, and to assess why different intervals yield such different paleolatitude estimates within a short window of geologic time. We completed comprehensive paleomagnetic analyses, microscopic investigations, and geochronologic studies, which lead us to conclude the following:

1. The paleomagnetic directions of Chen et al. [2010] and Chen et al. [2014] from E<sub>1d</sub> lavas are reproducible.

2. Magnetic carriers for E<sub>1d</sub> are magmatic titanomagnetite and secondary hematite, whereas magnetic carriers for E<sub>2p</sub> are dominated by magmatic titanomagnetite with negligible amounts of secondary hematite.

3. Whole rock <sup>40</sup>Ar/<sup>39</sup>Ar ages of lavas from E<sub>1d</sub> are ~10 Myr younger than zircon U-Pb ages, suggesting a post-emplacement thermal disturbance. In contrast, whole rock <sup>40</sup>Ar/<sup>39</sup>Ar ages of lavas from E<sub>2p</sub> are consistent with zircon U-Pb ages.

4. Volcanic rocks from E<sub>1d</sub> are probably remagnetized due to acquisition of secondary chemical remanent magnetization carried by post-emplacement hematite and thermoviscous remanent magnetization carried by titanomagnetite.

5. Volcanic rocks from E<sub>1d</sub> cannot be used to determine the Paleogene paleolatitude of the Lhasa terrane. All available evidence, however, suggest that the Paleogene paleolatitude of the Lhasa terrane of ~20 ± 4°N determined from volcanic rocks from E<sub>2p</sub> is reliable. Our results reported here, as well as those published elsewhere, suggest that the Tibetan Himalaya-Lhasa collision was underway by or slightly before 52 Ma.

6. In addition to a secondary origin of hematite as cause for the remagnetization, we also
detect the acquisition of secondary thermoviscous remanent magnetization in the volcanic rocks from the lower Linzizong Group. Future paleomagnetic studies in tectonically active regions should be vigilant in assessing sampled units for hydrothermal fluid activity with related thermal anomalies.

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Figure 1. (a) Simplified geological map showing the distribution of the Gangdese batholith and associated Linzizong volcanic successions in the Lhasa terrane, southern Tibet (modified from Yin and Harrison [2000] and Lee et al. [2009]); (b) Simplified geologic map of the Linzhou basin (modified from He et al. [2007]). Diamonds and dots represent paleomagnetic sampling locations: white diamond- Dupont-Nivet et al. [2010b] from lower E<sub>2p</sub>, green diamond- Huang et al. [2013] from upper E<sub>2p</sub>, blue diamond- this research from E<sub>1d</sub>. Red dot: Chen et al. [2010] and Chen et al. [2014] from E<sub>1d</sub>. Red stars represent geochronological sampling locations.
Figure 2. Field observations of the lower part of the Linzizong Group in the Linzhou basin. (a, b) Lava flows at the bottom of E₁d are fresh with minor surficial alteration or alteration along cracks. (c, d) Lava flows in the middle part of the sampled E₁d are altered to yellow-red. (e) Sampled lava flows from E₁d: note the upsection change in color from blue-gray to red. (f) All lava flows in the upper part of E₁d are red colored. (g) Alteration of the lava flows close to dykes is characterized by precipitation of secondary hematite in microfractures. (h) E₁d is pervasively intruded by dykes. (i) Strongly jointed lava flows from the top of E₁d. (j, k) Dykes cut across the unconformity between E₁d and E₂n. (l) Strata of E₂n above the unconformity are composed of red mudstone, conglomerate, and limestone.
Figure 3. (a-c) $^{40}$Ar/$^{39}$Ar and U-Pb zircon dating of the volcanic samples collected from both $E_{2p}$ (GS1-4 and GS124) and $E_{1d}$ (SH01) of the Linzizong Group in the Linzhou basin. (d) Cumulative age probability plot of dating results of SH01; a single zircon crystal gives concordant Neogene age in addition to the results. (e) Individual zircon analyses of LD8 from top of the $E_{1d}$. (f) Individual zircon analyses of LD9 above the Cretaceous redbeds yield Neogene ages.
Figure 4. Demagnetization diagrams for samples representing the entire E1d interval. All diagrams are displayed after bedding tilt correction. Closed (open) symbols represent the projection of vector end-points on the horizontal (vertical) plane; values represent alternating field and thermal demagnetization steps in milliTesla (mT) and °C, respectively. Most of the
samples only yielded interpretable results when processed with thermal demagnetization (a-j and i-r); AF demagnetization can only remove part of the low-temperature component (e.g., SH16.1 in k). Few samples (SH19.4, m) yield comparable characteristic remanent magnetizations by AF and thermal treatments.
Figure 5. (a) Equal-area projections of the isolated low-temperature component in geographic coordinates. Red dots (circles) are rejected by the 45° cut-off criterion. (b, c) Equal-area projections of site-mean ChRM directions before (IS, b (‘geographic’ on the figure itself)) and after (TC, c (‘stratigraphic’ on the figure itself)) bedding tilt correction for volcanic sites from E1d. Blue circles are the results from volcanic rocks from E1d reported by Chen et al. [2010] and Chen et al. [2014]. The common true mean direction test of the mean directions from the two studies after tilt correction is positive with classification of “C” [McFadden and Lowes, 1981]. Star represents present day magnetic field (PDF) at the sample locality.
Figure 6. Magnetic anisotropy of E$_{1d}$ volcanic rocks from AMS measurements. (a) Lineation (L) versus Foliation (F) diagram. (b) P versus T plot. $P = K_{\text{max}}/K_{\text{min}}$, $F = K_{\text{int}}/K_{\text{min}}$, $L = K_{\text{max}}/K_{\text{int}}$, $T = (\ln F - \ln L)/(\ln F + \ln L)$ [Jelinek, 1981]. (c, d) Stereoplots (equal area, lower-hemisphere projection) showing AMS data of E$_{1d}$ volcanic rocks. Maximum, intermediate, and minimum principal anisotropy axes are indicated, respectively, by triangles, squares and circles. Open symbols are mean directions, with 95 percent confidence zones indicated. TC, tilt correction; IS, in situ (without tilt correction).
Figure 7. (a-c) High-field thermomagnetic runs on a modified horizontal translation Curie balance for typical samples from E_{1d}. Thin (thick) lines represent the heating (cooling) curves, which are also indicated by the arrows (heating: red, and cooling: blue). (d-f) Hysteresis loops,
IRM, and back-field curves for characteristic samples from E1d. Hysteresis loops are corrected for the paramagnetic contribution. The sample codes are indicated at the top of each diagram and the hysteresis parameters are provided in the bottom-right part of the panels (\(M_s\): saturation magnetization, \(M_r\): remanent saturation magnetization, and \(B_c\): coercive force). (g-i) First-order-reversal-curve (FORC) diagrams for typical samples processed with FORCinel 2.0 [Harrison and Feinberg, 2008]. The applied smoothing factor (SF) is marked in the top-right part of each diagram. (j-l) Representative examples of the IRM component analysis [Kruiver et al., 2001]. Four components (Components L1-L4) are required to fit the IRM curves. Squares are measured data points. The components are marked with different lines: the linear acquisition plot (LAP) and the gradient acquisition plot (GAP) are shown in hatches.
Figure 8. End member modeling for volcanic rocks from E1d. (a) Diagram showing the $r^2$ versus number of end-members. (b, c) End-members for the normalized IRM acquisition curves in the three (b) and four (c) end-member models. (d) Ternary plot of the three end-member model showing the end-member percentages for individual specimens. (e, f and g) IRM component analysis [Kruiver et al., 2001] of the end-members in the preferred three
end-member model. Squares are the calculated data points from end-member modeling that are used to fit the IRM components. The components are marked with different lines: the linear acquisition plot (LAP) and the gradient acquisition plot (GAP) are shown. Component D1 has a very low coercivity and contribution; it is not assigned a physical meaning. Components D2 and D3 are harder; they represent Ti-poor and Ti-rich titanomagnetite respectively. Component D4 has the highest coercivity, and shows properties characteristic of hematite. IRM is normalized to the highest IRM-value (IRM/IRM$_{700}$); log$_{10}$ (B$_{1/2}$) and DP are in log$_{10}$ mT. Values of B$_{1/2}$ are displayed on each panel.
Figure 9. End member modeling for volcanic rocks from E$_2$p. (a) Diagram showing the $r^2$ values versus the number of end-members used to model the IRM. (b, c) End-members for the normalized IRM acquisition curves in the three (b) and four (c) end-member models. (d) Ternary plot showing percentages of the three end-members in the end-members for each specimen. (e, f, and g) IRM component analysis [Kruiver et al., 2001] of the end-members in
the preferred three end-member model. Squares are the calculated data points from end-member modeling that are used to fit IRM components. The components are marked with different lines: the linear acquisition plot (LAP) and the gradient acquisition plot (GAP) are shown. Component P1 has a very low coercivity and contribution; it is not assigned a physical meaning. Components P2 and P3 are harder; they represent Ti-poor and Ti-rich titanomagnetite respectively. Component P4 has the highest coercivity, and shows properties characteristic of hematite. IRM is normalized to the highest IRM-value (IRM/IRM_{700mT}); log_{10} (B_{1/2}) and DP are in log_{10} mT. Values of B_{1/2} are displayed on each panel.
Figure 10. Thin section photographs of representative volcanic samples from both E1d and E2p under plane polarized light. (a-c) SH2.8 and SH5.5 from the lower part of E1d, showing a porphyritic texture and alteration of feldspar lapillls and silicic matrix. (d) Precipitation of secondary hematite along the a crack in the rock, with abundance decreasing inwardly in sample SH18.4 from the upper part of E1d. (e, f) Secondary hematite is distributed along the cracks of the minerals in samples SH18.6 and SH18.7. (g) Ignimbrite sample GL2.3 from E2p, showing a pyroclastic texture with scattered rock lapillls distributed in an ash matrix. (h, i) Ignimbrite samples GL9.7 (h) and GL26.8 (i) from E2p with porphyritic texture show little sign of alteration.
Figure 11. SEM back-scattered electron images for selected samples from both $E_{1d}$ and $E_{2p}$.

Hem: hematite, Ti-Mt: titanomagnetite. (a, b and c) Samples SH2.4 and SH4.2 from the bottom of $E_{1d}$, with abundant titanomagnetite particles showing solid-state exsolution features. (d-f) Hematite growing around or within the silicate minerals in sample SH13.3 from the middle part
(g) High-magnification view of sample SH13.3 showing secondary hematite growth around a large titanomagnetite grain. (h, i) Small titanomagnetite particles with low abundance disperse in sample SH18.4, whereas fine hematite particles are everywhere. Also observed is hematite replacing a titanomagnetite grain with the original crystal habit of titanomagnetite retained in (h). (j, k) Alteration of titanomagnetite to fine-grained hematite in samples SH24.8 from the top of E1d. (l) Euhedral titanomagnetite grains with typical solid-state exsolution features in sample SH24.8. (m) Titanomagnetite with variable grain sizes in the rock lapillis from ignimbrite sample GL8.5 from E2p. (n) Large titanomagnetite grain with solid-state exsolution features in sample GL27.3 from ignimbrite sample GL8.5 from E2p.
Figure 12. Possible sequence of events for the remagnetization of \( E_{1d} \) in the Linzhou basin in cross section view. (a) Eruption of \( E_{1d} \) at \( \sim 64-62 \) Ma (Chron C26r, or C27r) at \( \sim 20^\circ \)N latitude; a primary remanent magnetization (PRM, solid black) is acquired. (b) Southward tilting of \( \sim 5^\circ \) of \( E_{1d} \) between \( \sim 62-54 \) Ma, remagnetization induced by paleoweathering happened and a chemical remanent magnetization (CRM\(_1\), dashed gray) carried by secondary hematite was acquired. \( E_{2n} \) was deposited subsequently at \( \sim 62-54 \) Ma. (c) Continued southward tilting of \( E_{1d} \) and \( E_{2n} \) to \( \sim 15^\circ \) before eruption of \( E_{2p} \) volcanic rocks. Subsequently, \( E_{2p} \) was erupted/deposited and massive dikes (pink) intruded \( E_{1d} \) and \( E_{2n} \) at \( \sim 54-48 \) Ma at \( \sim 20^\circ \)N.
latitude. Strong remagnetizations were induced due to thermochemical and thermoviscous overprinting. Hematite formed during the high-temperature hydrothermal alteration and a second CRM (CRM$_2$, dashed gray) was acquired. Combination of CRM$_1$ and CRM$_2$ results in a remanent magnetization (solid red) carried by hematite with low inclination. Part of Ti-poor titanomagnetite acquired a thermoviscous remanent magnetization (TVRM$_1$, dashed gray). A combination of PRM and TVRM$_1$ resulted in a remanent magnetization (solid blue) carried by Ti-poor titanomagnetite with low inclination. PRM carried by Ti-rich titanomagnetite is completely reset to TVRM$_1$. (d) Northward tilting of ~30° of the entire sequence between 48 Ma and 13 Ma. (e) Remanent magnetization carried by Ti-rich titanomagnetite was erased again after tilting and a new TVRM (TVRM$_2$, dashed purple) was acquired at ~20°N latitude at ~13 Ma related to Neogene dikes (green) intruded into E$_{1d}$. Note that E$_{1d}$ and E$_{2n}$ are ~270 m and ~330 m thick, respectively, whereas the volcanic rocks in E$_{2p}$ have a thickness of ~1900 m [He et al., 2007].